



Please write clearly in block capitals.

Centre number

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Candidate number

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Surname

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Forename(s)

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Candidate signature

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# A-level PHYSICS A

Unit 5A Astrophysics  
Section B

Wednesday 21 June 2017

Morning

Time allowed: The total time for both sections of this paper is 1 hour 45 minutes. You are advised to spend approximately 50 minutes on this section.

## Materials

For this paper you must have:

- a calculator
- a pencil and a ruler
- a Data and Formulae Booklet (enclosed).

## Instructions

- Use black ink or black ball-point pen.
- Fill in the boxes at the top of this page.
- Answer **all** questions.
- You must answer the questions in the spaces provided.  
Do not write outside the box around each page or on blank pages.
- Do all rough work in this book.  
Cross through any work you do not want to be marked.
- Show all your working.

## Information

- The marks for questions are shown in brackets.
- The maximum mark for this section is 35.
- You are expected to use a calculator where appropriate.
- A *Data and Formulae Booklet* is provided as a loose insert.
- You will be marked on your ability to:
  - use good English
  - organise information clearly
  - use specialist vocabulary where appropriate.

For Examiner's Use	
Examiner's Initials	
Question	Mark
1	
2	
3	
4	
TOTAL	



J U N 1 7 P H Y A 5 2 A 0 1

WMP/Jun17/E6

PHYA5/2A

**Section B**

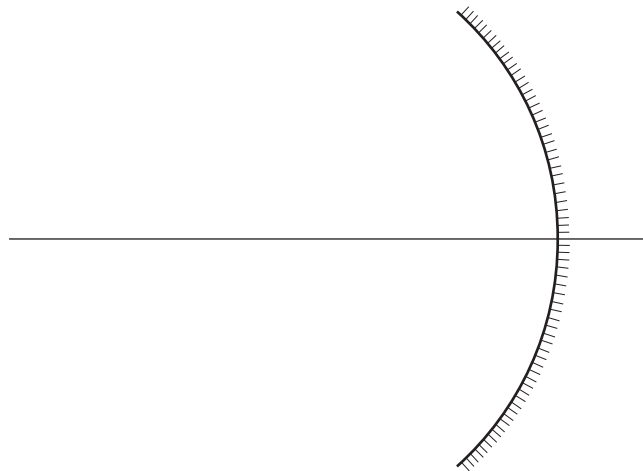
The maximum mark for this section is 35. You are advised to spend approximately 50 minutes on this section.

**1 (a)** Spherical aberration can be a problem with reflecting telescopes.

**1 (a) (i)** Complete the ray diagram in **Figure 1** to show how spherical aberration occurs in a reflecting telescope.

**[2 marks]**

**Figure 1**



**1 (a) (ii)** State how this problem can be prevented.

**[1 mark]**

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**1 (b)** A refracting telescope can produce a clearer image than a reflector of similar diameter because of the position of the secondary mirror.

**1 (b) (i)** Sketch a diagram to show the positions of the mirrors in a Cassegrain telescope.

**[1 mark]**



- 1 (b) (ii) Give **two** reasons why the secondary mirror in the Cassegrain telescope affects the clarity of the image.

[2 marks]

1 \_\_\_\_\_

\_\_\_\_\_

\_\_\_\_\_

2 \_\_\_\_\_

\_\_\_\_\_

\_\_\_\_\_

- 1 (c) The Very Large Telescope (VLT) in the Atacama Desert in Chile is a combination of four Cassegrain telescopes each of diameter 8.2 m. It is used to detect electromagnetic radiation of wavelengths in the range 300 nm to 20  $\mu\text{m}$ .

- 1 (c) (i) Show that the combination has a similar light-collecting power to that of a single telescope of diameter 16 m.

[2 marks]

Question 1 continues on the next page

Turn over ►



- 1 (c) (ii) The VLT is capable of an angular resolution similar to that of a 130 m diameter telescope.

Calculate the minimum angular resolution of the VLT.

[1 mark]

minimum angular resolution = \_\_\_\_\_ rad

- 1 (c) (iii) What part of the electromagnetic spectrum is significantly absorbed by water vapour?  
Tick (✓) **one** box next to the correct answer.

[1 mark]

Infrared

Radio waves

Ultraviolet

X-rays

10



**Turn over for the next question**

**DO NOT WRITE ON THIS PAGE  
ANSWER IN THE SPACES PROVIDED**

**Turn over ►**



2 **Table 1** gives some properties of the Garnet Star.

**Table 1**

apparent magnitude	4.08
absolute magnitude	− 7.63
surface temperature	3750 K

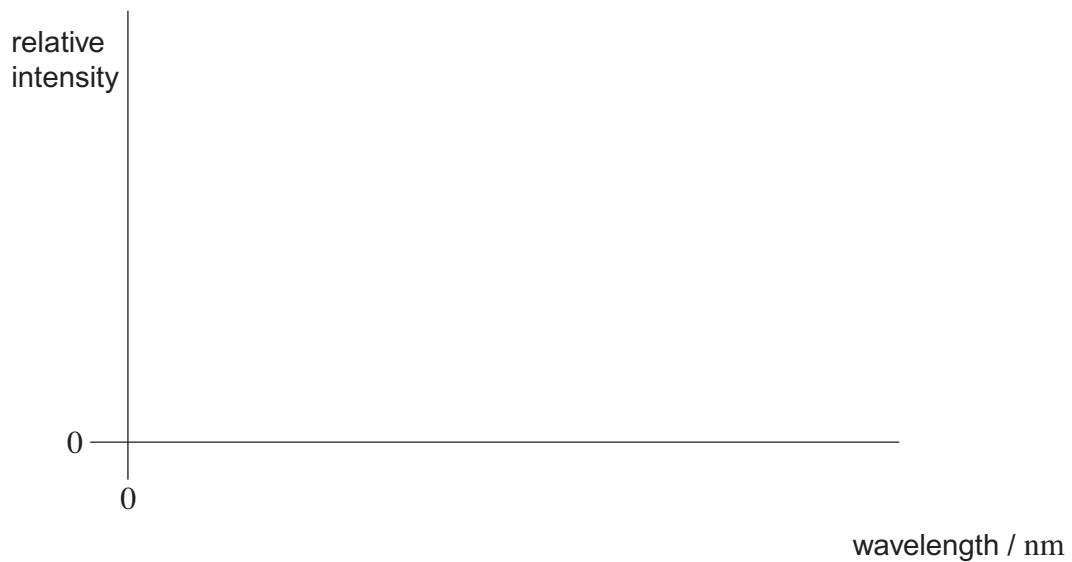
2 (a) (i) Calculate the wavelength of the peak in the black-body radiation curve for the Garnet Star.

[1 mark]

wavelength = \_\_\_\_\_ m

2 (a) (ii) Sketch the black-body radiation curve for the Garnet Star on the axes below. Label the wavelength axis with a suitable scale.

[2 marks]



**2 (b)** Calculate, in ly, the distance from Earth to the Garnet Star.

**[3 marks]**

distance = \_\_\_\_\_ ly

**2 (c)** The Garnet Star is one of the largest stars yet discovered. Its radius is  $1.2 \times 10^9$  km.

Calculate the ratio  $\frac{\text{power output of the Garnet Star}}{\text{power output of the Sun}}$ .

surface temperature of the Sun = 5800 K  
radius of the Sun =  $6.9 \times 10^5$  km

**[2 marks]**

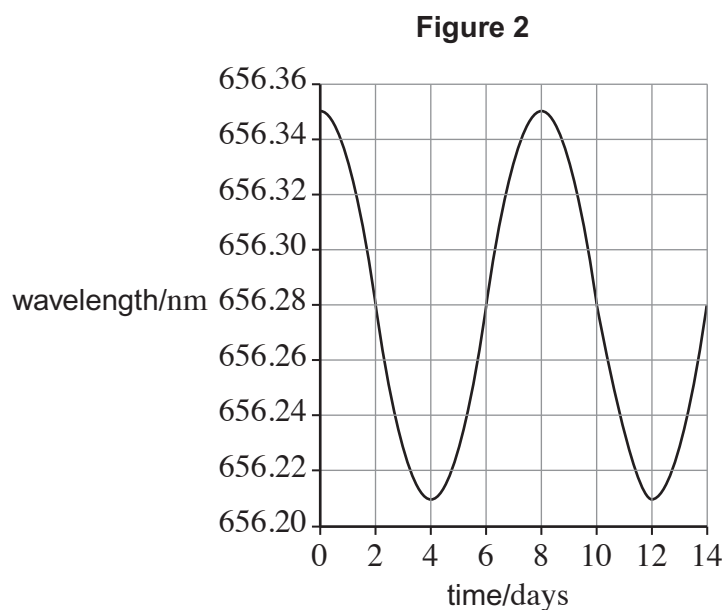
ratio = \_\_\_\_\_

8

Turn over ►



- 3 Eta Orionis is an eclipsing binary system. Analysis of the light from one of the stars shows that a particular spectral line varies in wavelength as shown in **Figure 2**.



- 3 (a) (i) Show that the star has an orbital speed of approximately  $30 \text{ km s}^{-1}$ .

[2 marks]

- 3 (a) (ii) Calculate the diameter of the orbit of the star.

[2 marks]

diameter = \_\_\_\_\_ m











4 (b) Measurements of the shift in the 21 cm line in the hydrogen spectrum of galaxy M84 suggest that it is receding at a velocity of  $1100 \text{ km s}^{-1}$ .

4 (b) (i) Calculate the value of the red shift  $z$  for this galaxy.

[1 mark]

$z =$  \_\_\_\_\_

4 (b) (ii) Calculate, in Mpc, the distance from Earth to this galaxy.

[2 marks]

distance = \_\_\_\_\_ Mpc

**END OF QUESTIONS**

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